

Simulation of Magnetic Spindown in Rotating Hot-Star Winds

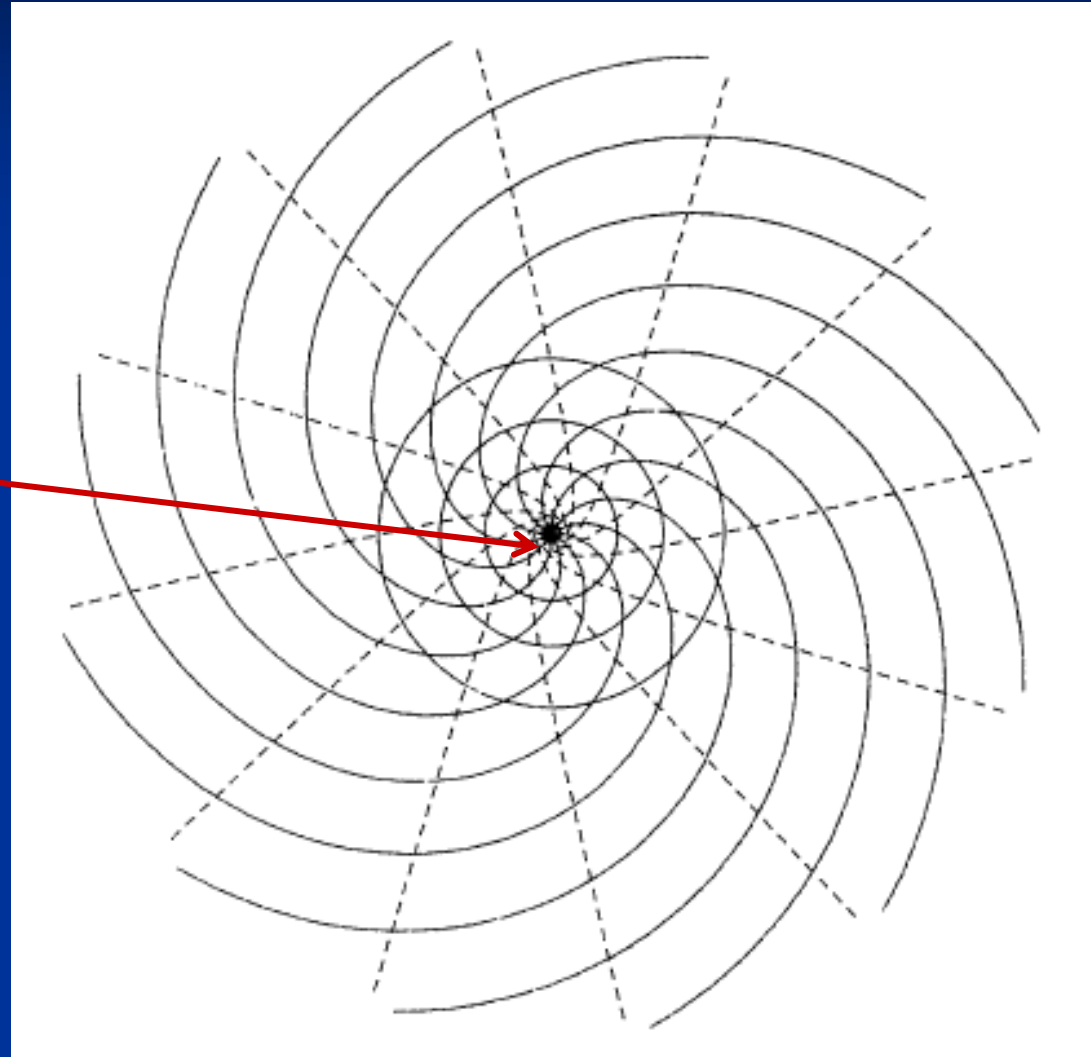
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Collaborators: Stan Owocki, Rich Townsend

Weber and Davis (1967)

Monopole field at
solar surface



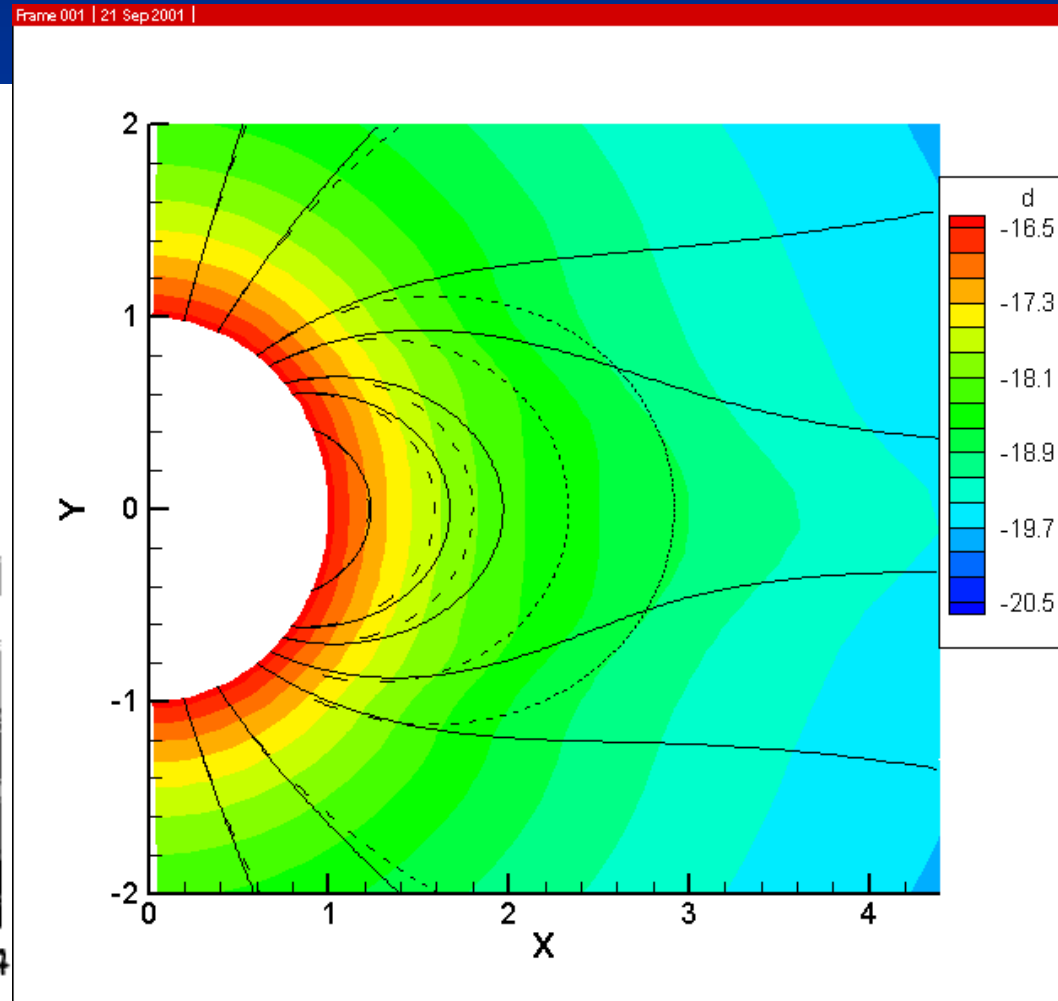
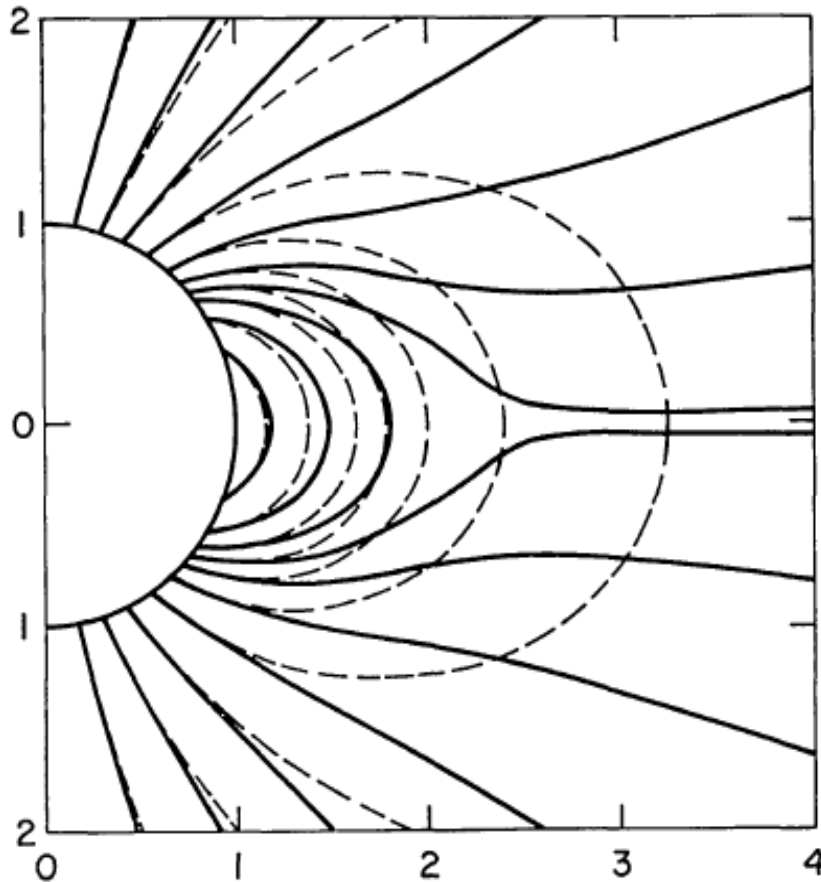
Pneuman and Kopp (1971)

Iterative scheme

Fully dynamic, **time dependent**

**MHD model for base dipole
with $B_0=1$ G**

Our Simulation



Wind Magnetic Confinement Parameter

$$\eta(r) = \frac{B^2 r^2}{\dot{M} v} = \underbrace{\left[\frac{B_0^2 R_*^2}{\dot{M} v_\infty} \right]}_{\equiv \eta_*} \left[\frac{(r/R_*)^{2-2q}}{(1 - (R_*/r))^\beta} \right]$$

Angular Momentum Loss

Contribution from the field

$$\dot{J}_B = \int -\frac{B_r B_\phi(r, \theta)}{4\pi} r \sin \theta dA$$

Contribution from gas

$$\dot{J}_{gas} = \int \rho v_r v_\phi r \sin \theta dA$$

Need to
compute

Total loss

$$\dot{J}_{tot} = \dot{J}_B + \dot{J}_{gas}$$

Weber and Davis

$$\dot{J}_{tot} \approx \frac{2}{3} \dot{M} \Omega R_A^2$$

As if gas co-rotates to R_A

Computing Alfven Radius

$$\eta(r) = \eta_* \left[\frac{(r / R_*)^{2-2q}}{(1 - (R_* / r))^\beta} \right] \quad \text{For hot stars, } \beta \sim 1$$

Alfven radius $\eta(R_A) \equiv 1$

For a monopole $q=2$

$$\frac{R_A}{R_*} \approx \eta_*^{1/2}$$

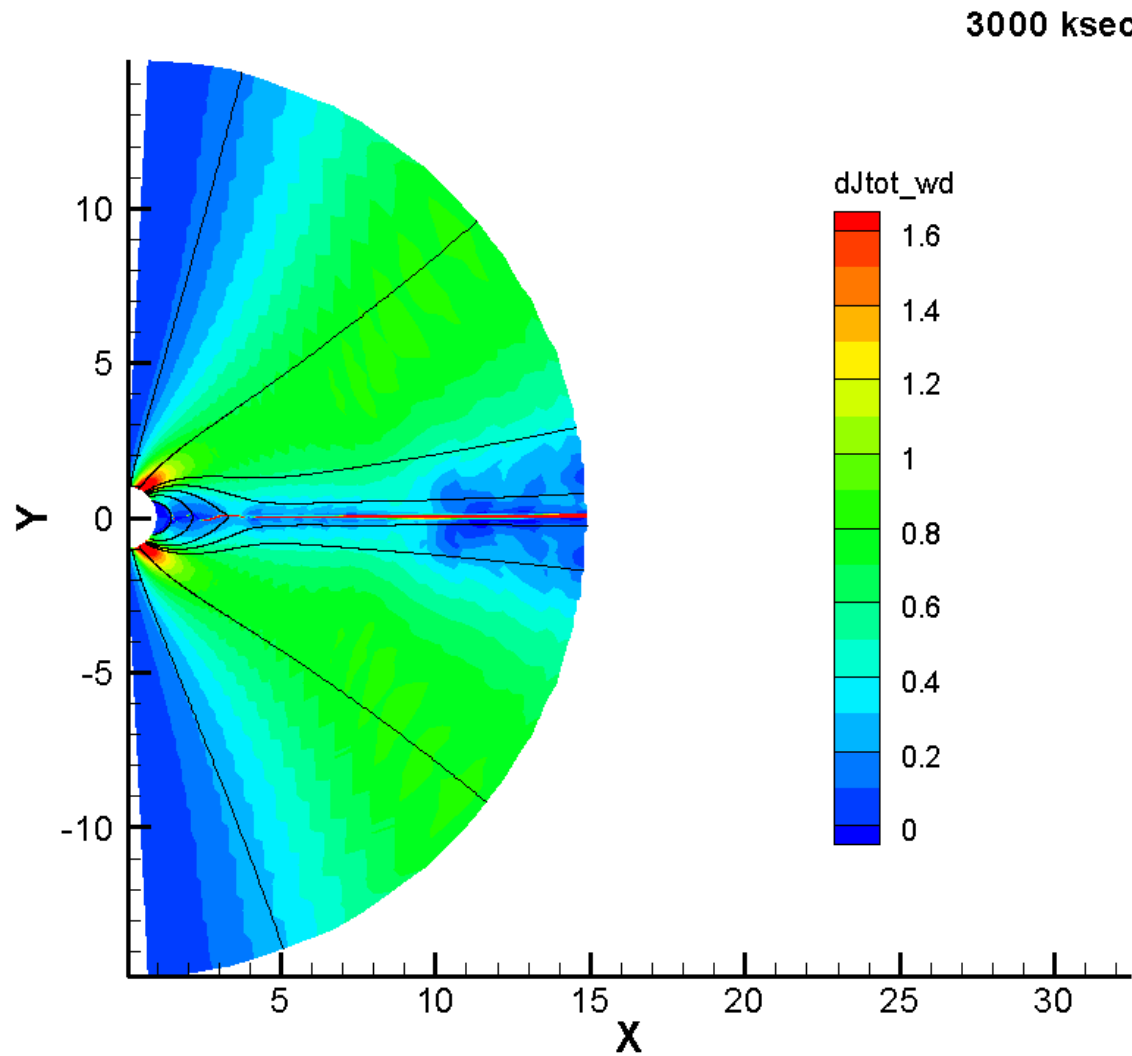
$$j \approx \frac{2}{3} \dot{M} \Omega R_*^2 \eta_*$$

For a dipole $q=3$

$$\frac{R_A}{R_*} \approx \eta_*^{1/4}$$

$$j \approx \frac{2}{3} \dot{M} \Omega R_*^2 \sqrt{\eta_*}$$

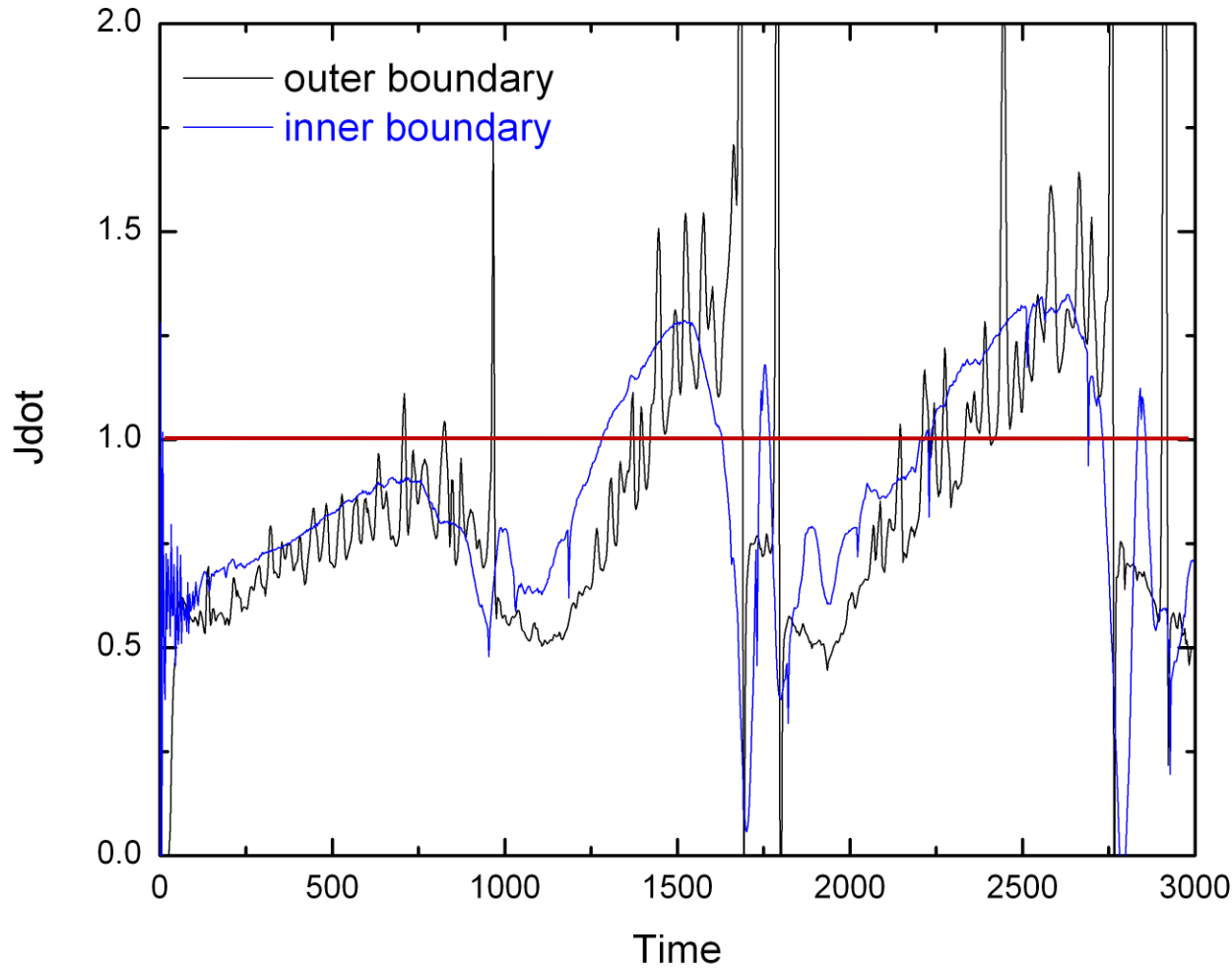
Angular Momentum Loss: Sims



Dead Zone Lives!

Angular Momentum Loss

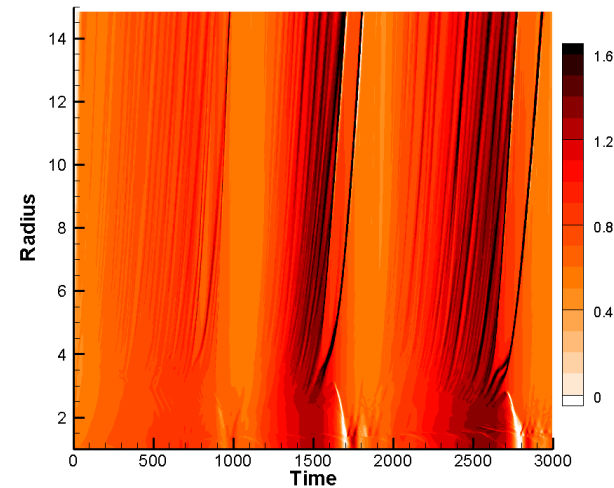
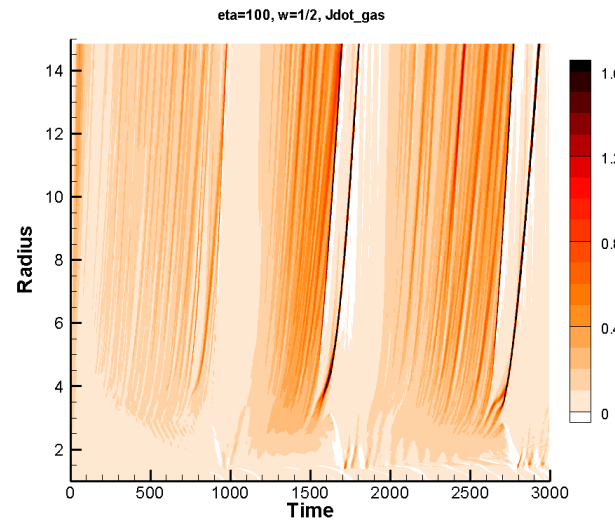
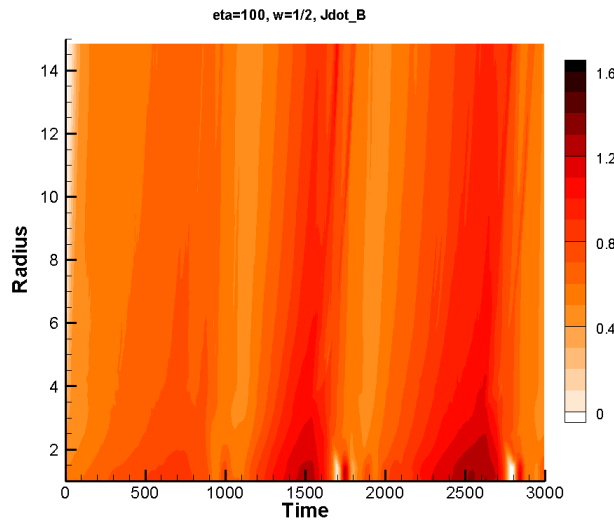
$\dot{J}/\dot{J}_{\text{WD Dipole}}$



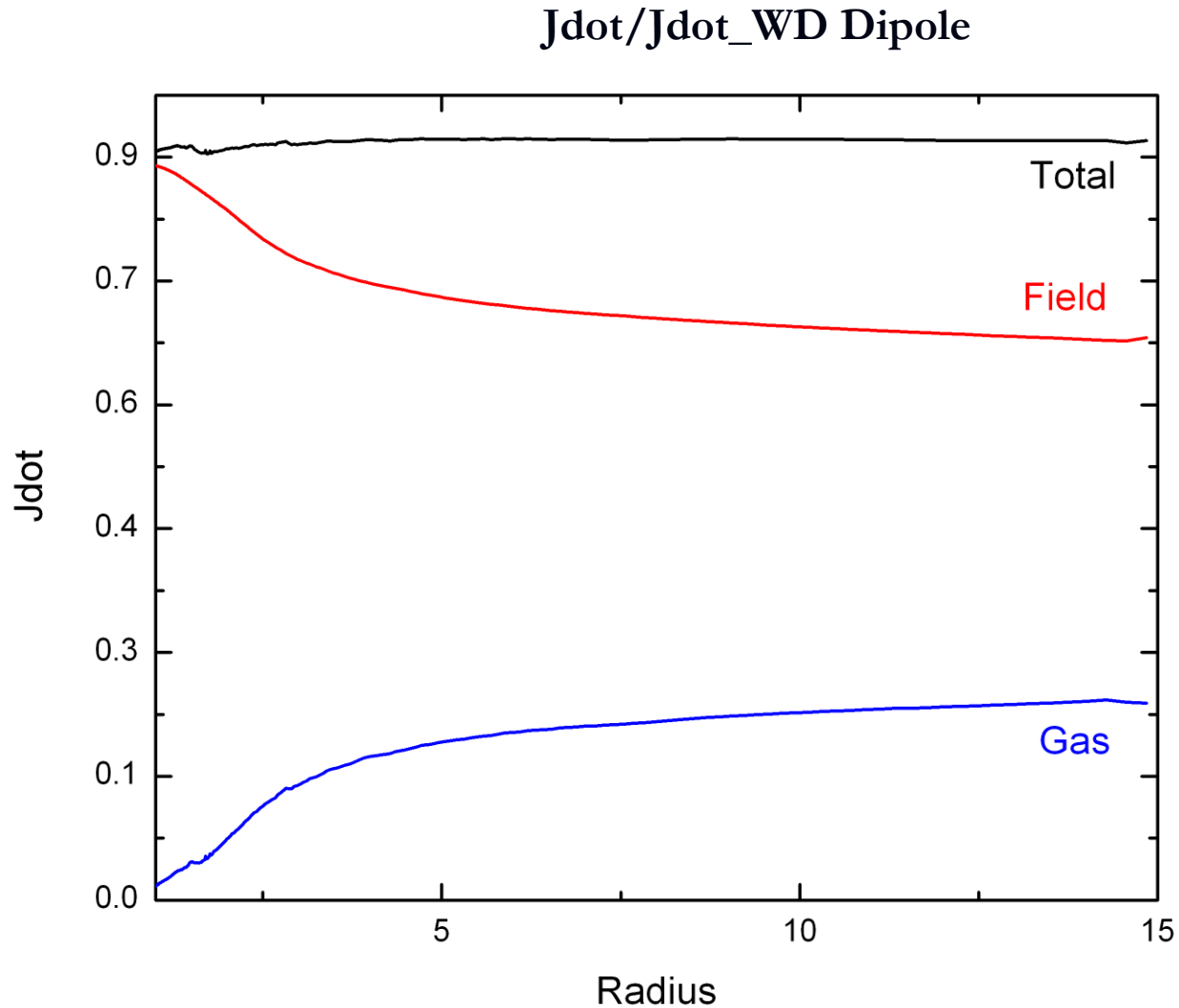
\dot{J} goes down during outbreaks.
Sorry, Trigiolo (2008)!

Angular Momentum Loss

Magnetic Field + Gas = Total



Time-Averaged Angular Momentum Loss



Angular Momentum Loss Weber and Davis

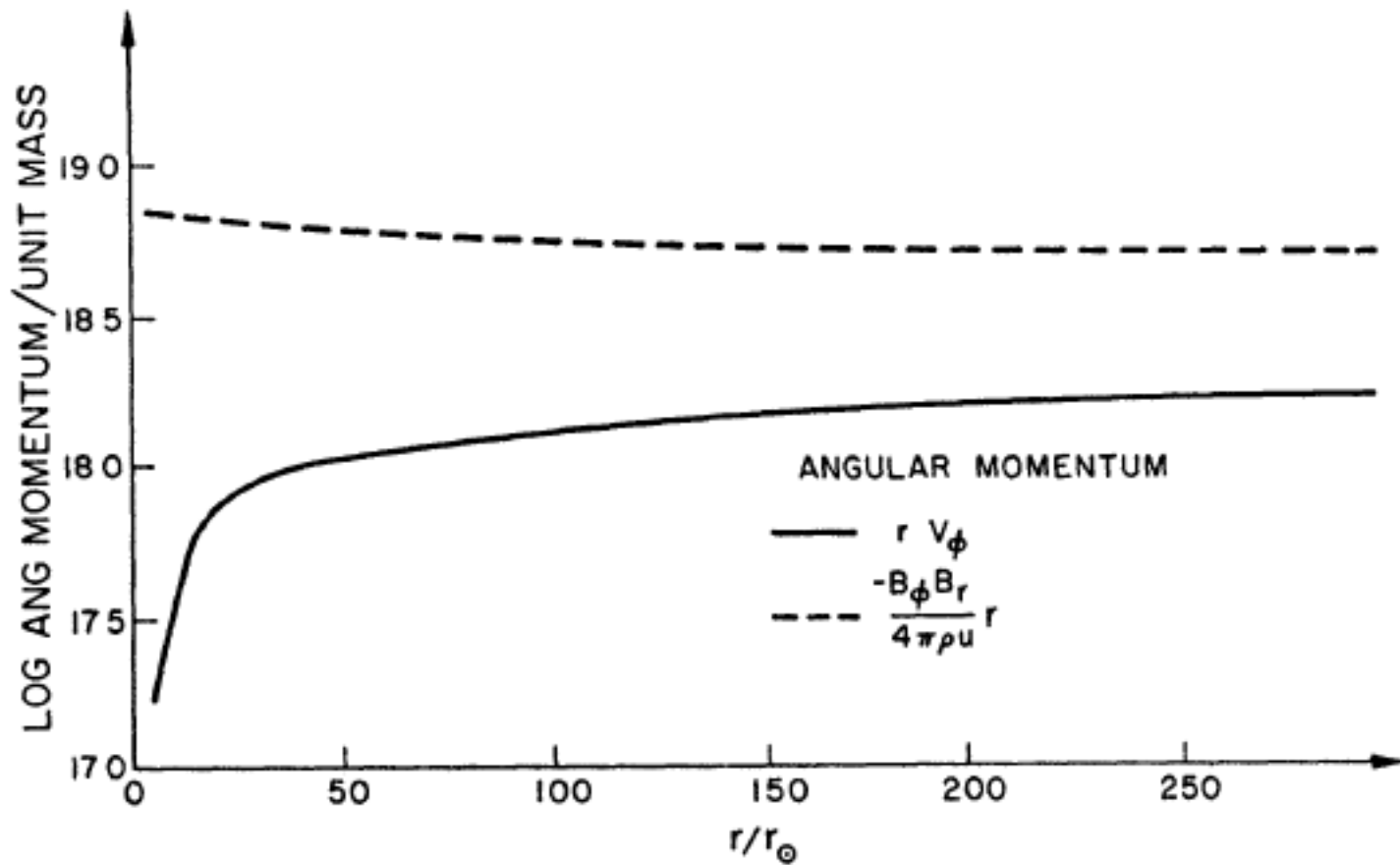


FIG. 5.—Angular momentum and magnetic torque in the solar wind

Spindown

$$\dot{J} \approx \frac{2}{3} \dot{M} \Omega R_A^2$$

contribution from both field and gas

For Dipole

$$\tau_{spin} \equiv \frac{J}{\dot{J}} \approx \frac{\frac{3}{2} I}{MR^2} \frac{M}{\dot{M}} \frac{1}{\sqrt{\eta_*}} = \tau_{mass} \frac{\frac{3}{2} k}{\sqrt{\eta_*}}$$

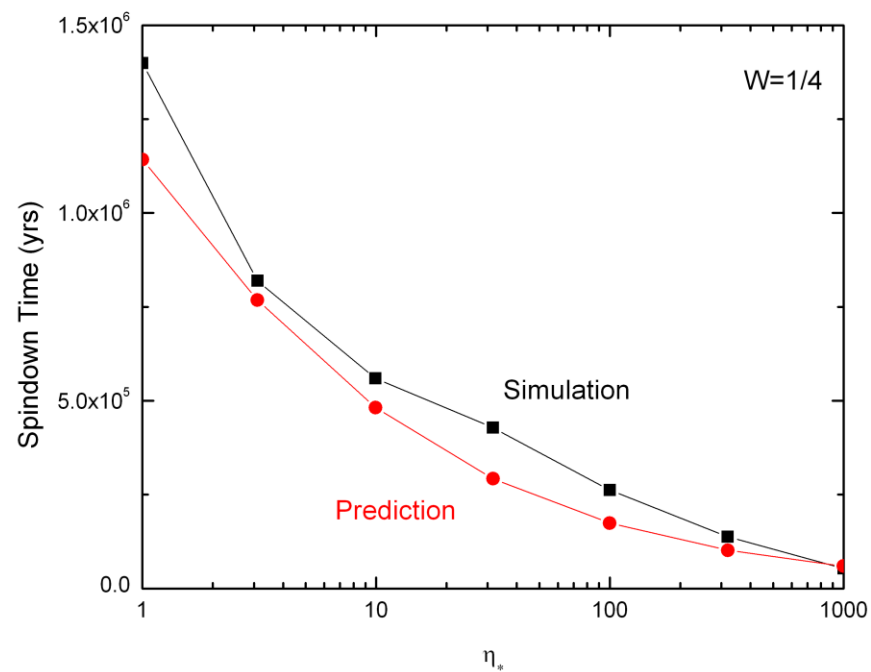
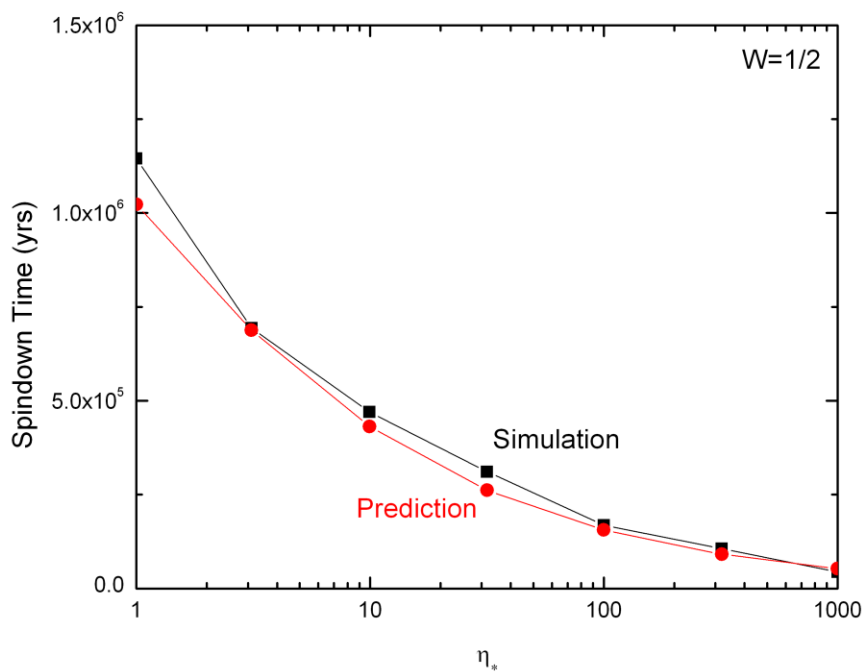
$$\frac{\tau_{spin}}{\tau_{mass}} \approx \frac{0.15}{\sqrt{\eta_*}}$$

Quiz: what is the scaling for monopole?

Spindown Time

$W=1/2$

$W=1/4$



Sample Spindown Time

Selected Stars on our target list

Name	M(Msun)	Mdot (Msun/yr)	Bp (G)	R(Rsun)	Vinf (km/s)	η_*	τ_m (Myr)	τ_{sm} (Myr)	τ_{sd} (Myr)
Theta 1 Ori C	44	4.0E-07	1100	8.2	2500	15.7	110.0	1.1	4.2
HD 191612	40	6.1E-06	1500	16.7	2600	7.6	6.6	0.1	0.4
Zeta Cas	8	3.0E-10	340	5.9	800	3.2E+03	2.7E+04	1.2	70.3
Sigma Ori E	8.9	1.0E-10	9600	5.3	420	1.2E+07	8.9E+04	1.1E-03	3.9

Monopole

Dipole

Conclusions

- Angular momentum loss $\dot{J} \propto \Omega R_A^2$
- Spindown time \sim Mass time / $\sqrt{\eta_*}$
- $\dot{J}_B > \dot{J}_{gas}$
- breakout events: J stored and released